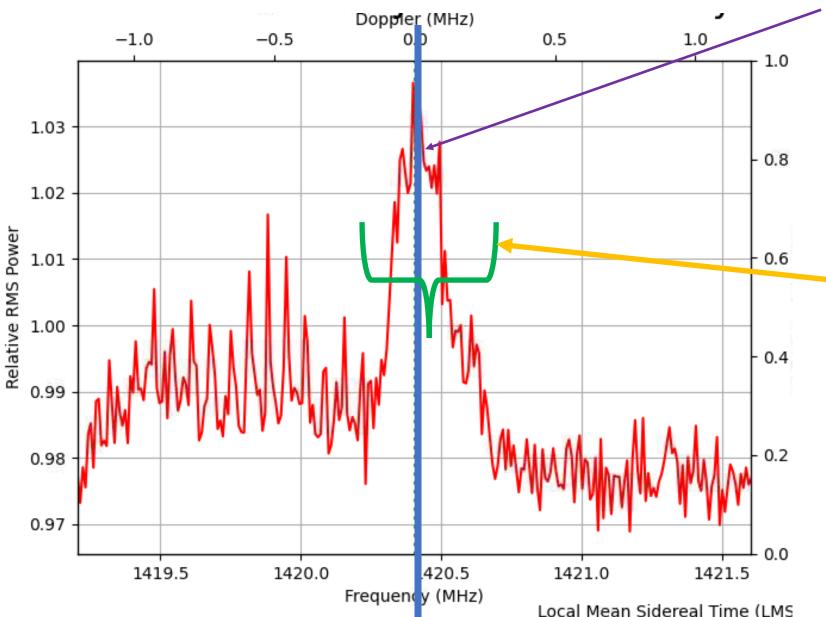


# Detecting molecular hydrogen line in Milky Way can now be achieved easily and cheaply

- Milky Way hydrogen detection used to be expensive but can now be achieved with very cheap equipment from Amazon/elsewhere, using Software Defined Radio (SDRs) and hydrogen line filter/pre-amp, and satellite dish/other aerial.
- The Society of Amateur Radio Astronomers (USA/"SARA") has produced a project called "Scope in a Box", which led me to give this a try.

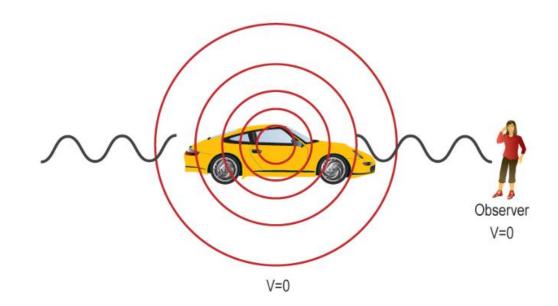
# Hydrogen line (1420.405MHz) Mapping Exercise of the Milky Way @ LRO

### An example of a detection of the hydrogen line

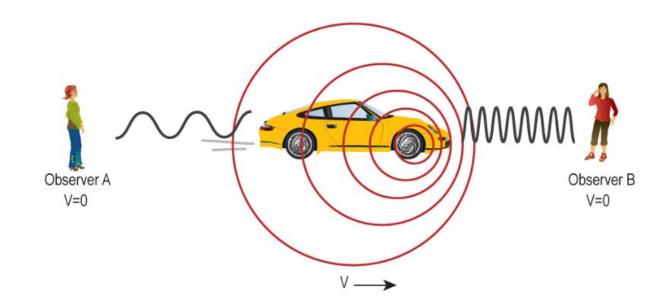


Frequency of hydrogen 1420.405MHz

Range of frequencies For hydrogen representing Doppler shift and hence differences in Velocity



### Doppler shift



#### Hydrogen Hydrogen Redshifted Hydrogen Blueshifted

Wavelength (nanometers)

# Red-shifted is moving away from us

### **Aerials & Mounts**



3m dishes come with mounting problems...

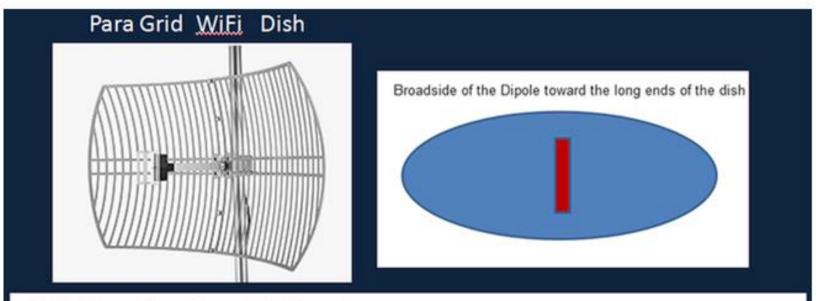


Corner aerials from chicken wire

Very big corner aerial

Perhaps more practical and easier to buy off the shelf is something Like this?

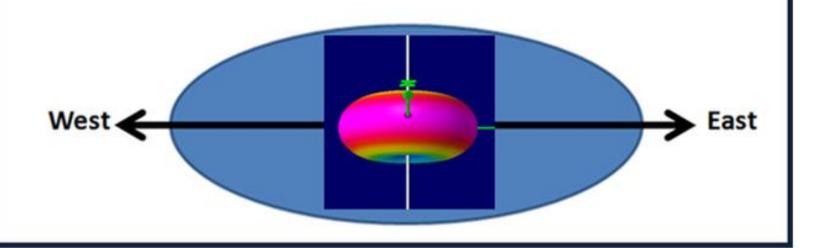




### WiFi 100cmx60cm Para-grid Orientation

Recall that beam-width is inversely proportional to dish size.

For optimal angular resolution of H1 drift scans, align the 1m side E<>W



From
Alex
Pettit

### Para Grid WiFi Dish

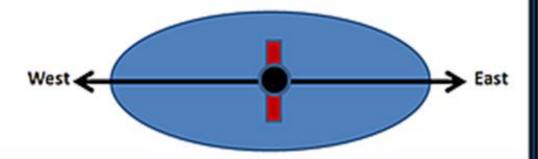
#### WiFi 100cmx50cm Para-grid Orientation

Recall that beam-width is inversely proportional to dish size.

1/2wl Dipole

1wl Loop

For optimal angular resolution of H1 drift scans, align the fm side E⇔W

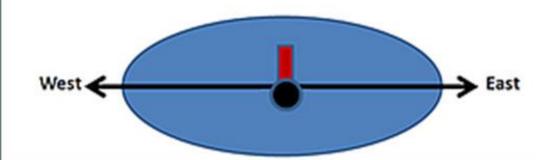


#### WiFi 100cmx60cm Para-grid Orientation

Recall that beam-width is inversely proportional to dish size.

1/4wl Mono

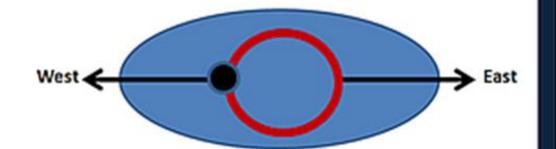
For optimal angular resolution of H1 drift scans, align the 1m side E<>W



#### WiFi 100cmx60cm Para-grid Orientation

Recall that beam-width is inversely proportional to dish size.

For optimal angular resolution of H1 drift scans, align the 1m side E⇔W



#### WiFi 100cmx60cm Para-grid Orientation

Recall that beam-width is inversely proportional to dish size.

For optimal angular resolution of H1 drift scans, align the 1m side E⇔W

West East

Patch From

Alex

Pettit



I have been very lucky & have been given one of these...

Ptarmigan Triffid
Military Phased Array
Not many of them
around!





# Nooelec SAWbird+ H1 - Saw Filter & Cascaded Ultra-Low Noise Amplifier (LNA)



# For folks starting to explore radio astronomy,

ezRA - Easy Radio Astronomy
Free 1420 MHz Galactic hydrogen
data collection and analysis

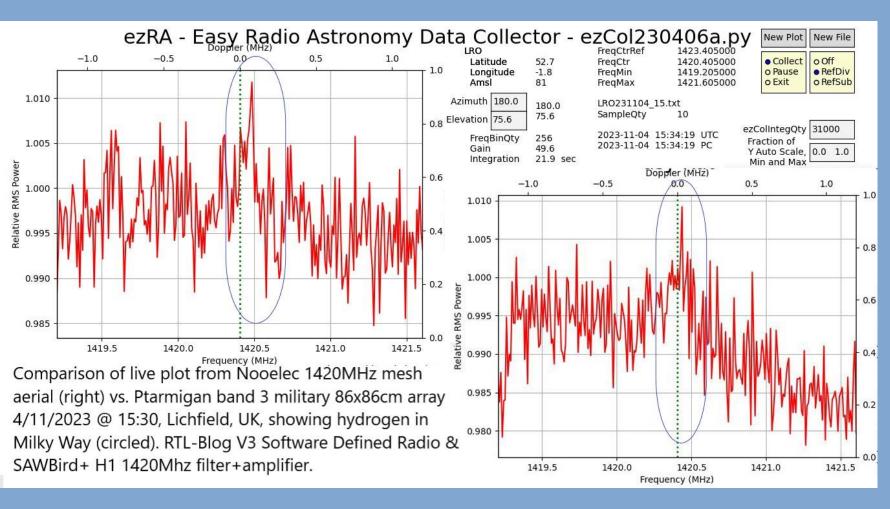
https://github.com/tedcline/ezRA

Windows and Linux

# ezRA software for collecting and processing data and mapping it over known background of radio sources in Milky Way

(Has own collection program, alternative= SDR# IFAverage Plug-In to collect data)

### Testing the new Nooelec 1420MHz-tuned mesh aerial





Quite capable of detecting hydrogen

## Data is improved if calibrated

Calibration of Scope in a Box in SDR# Software with IF Average plug in

This does similar job to flats and darks in astrophotography

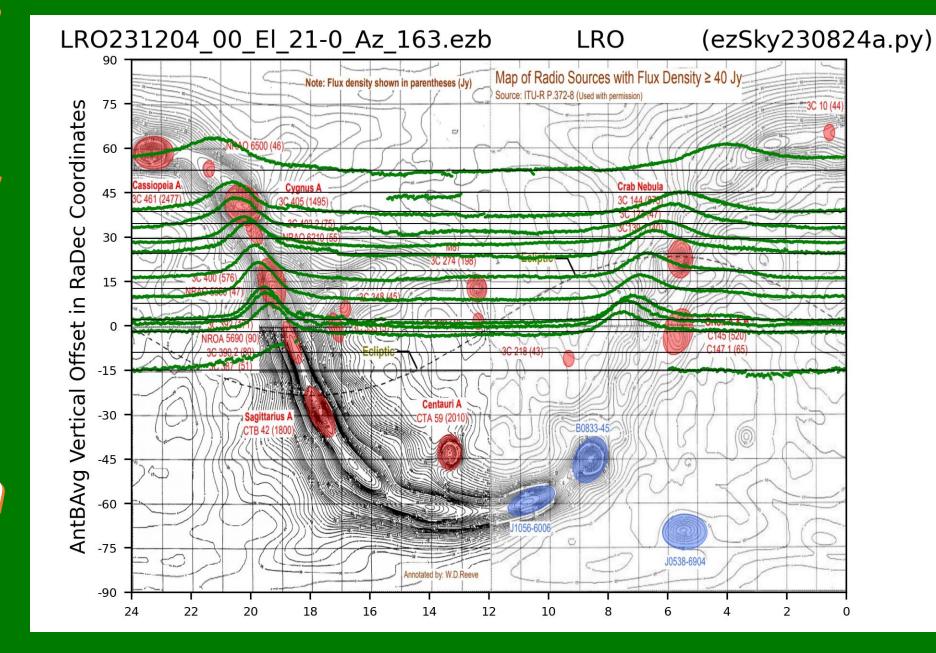
## 50 ohm load used in place of aerial



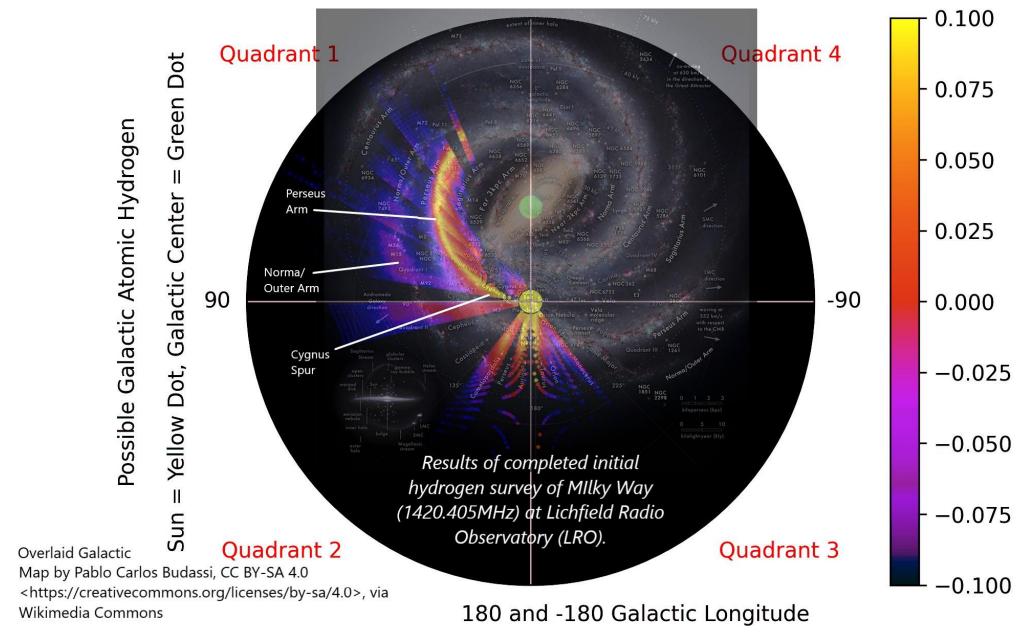


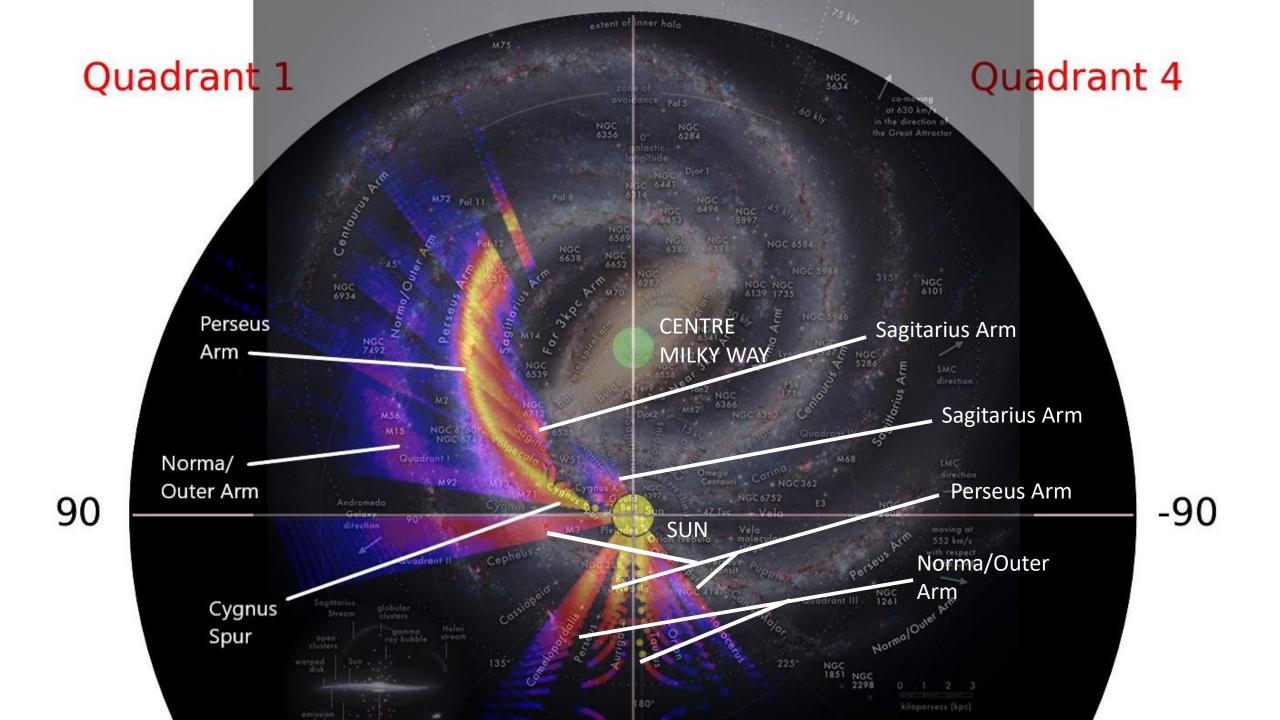
Although now I use signal offset 3MHz to 1423MHz every other sample for calibration.

Results of the first LRO Milky Way Hydrogen Map - elevation coverage

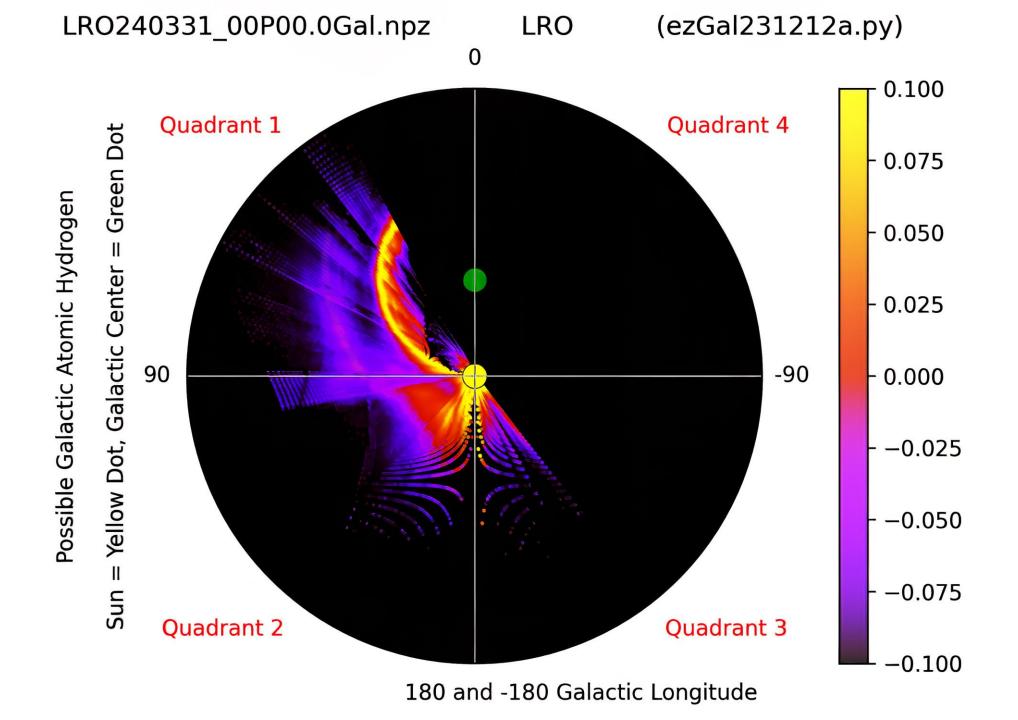


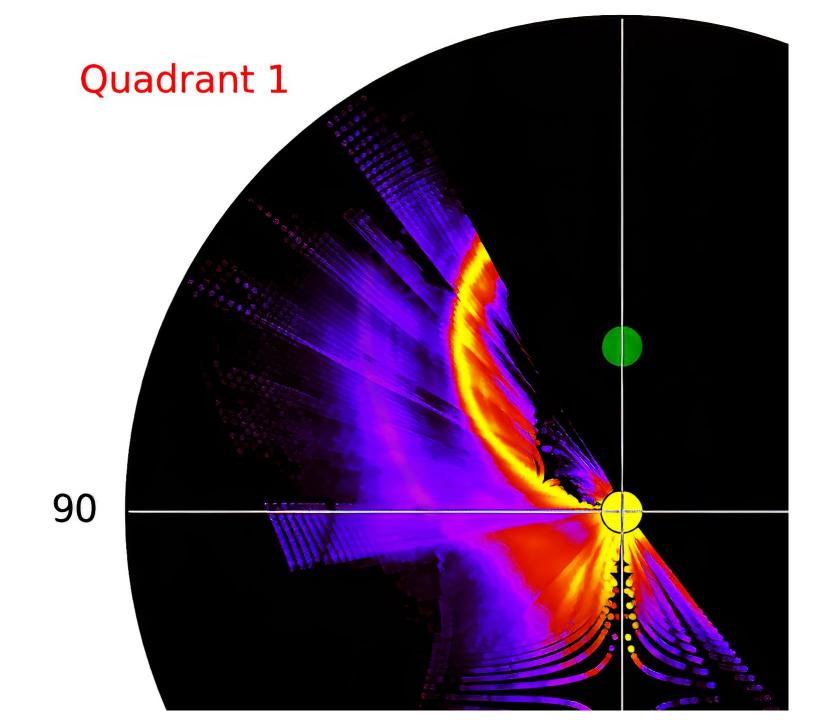
### Mapping the Milky Way Arms

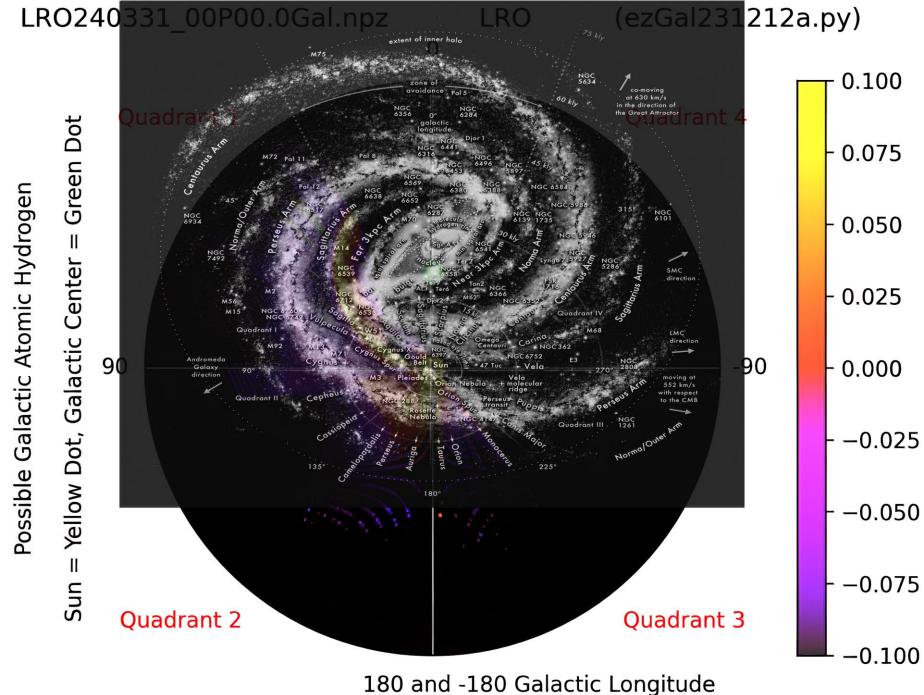




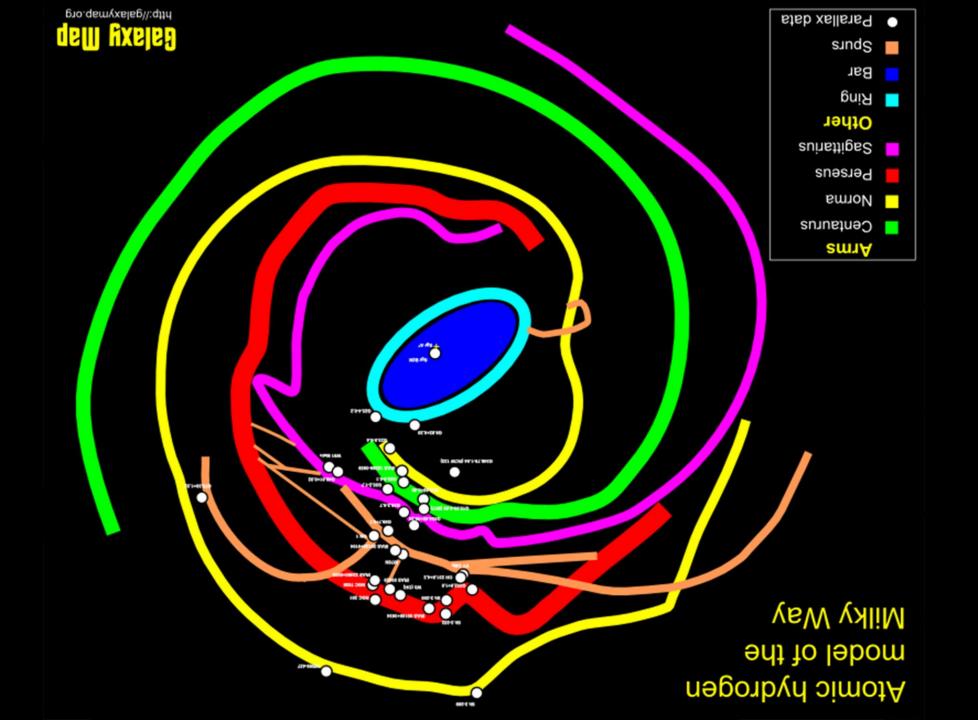
# Latest data from LRO 1/4/2024 coming next — note I now think I didn't match the two maps correctly in above diagrams

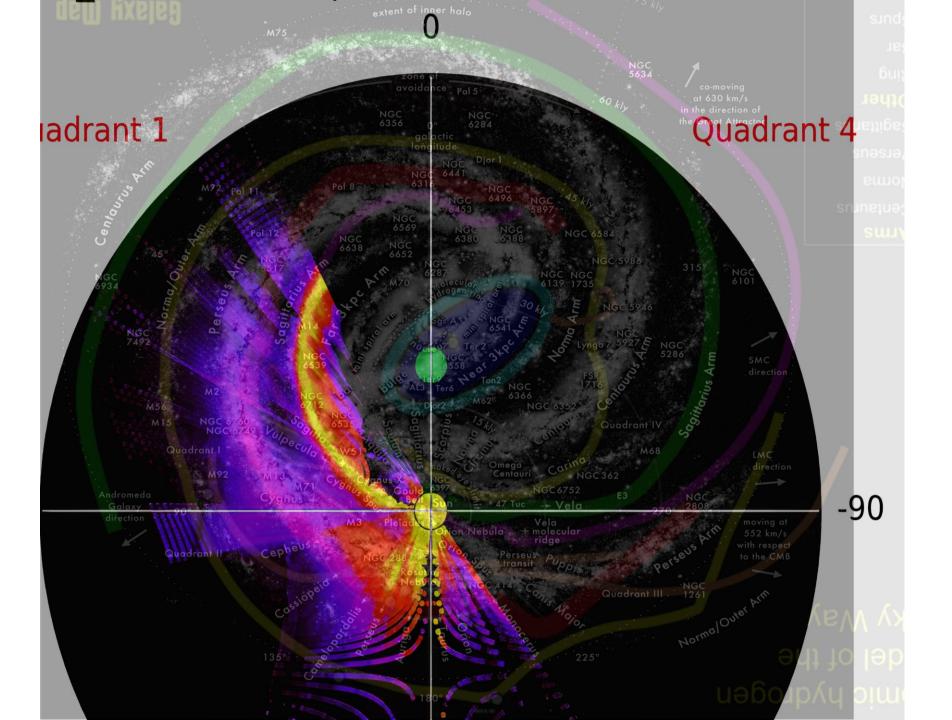


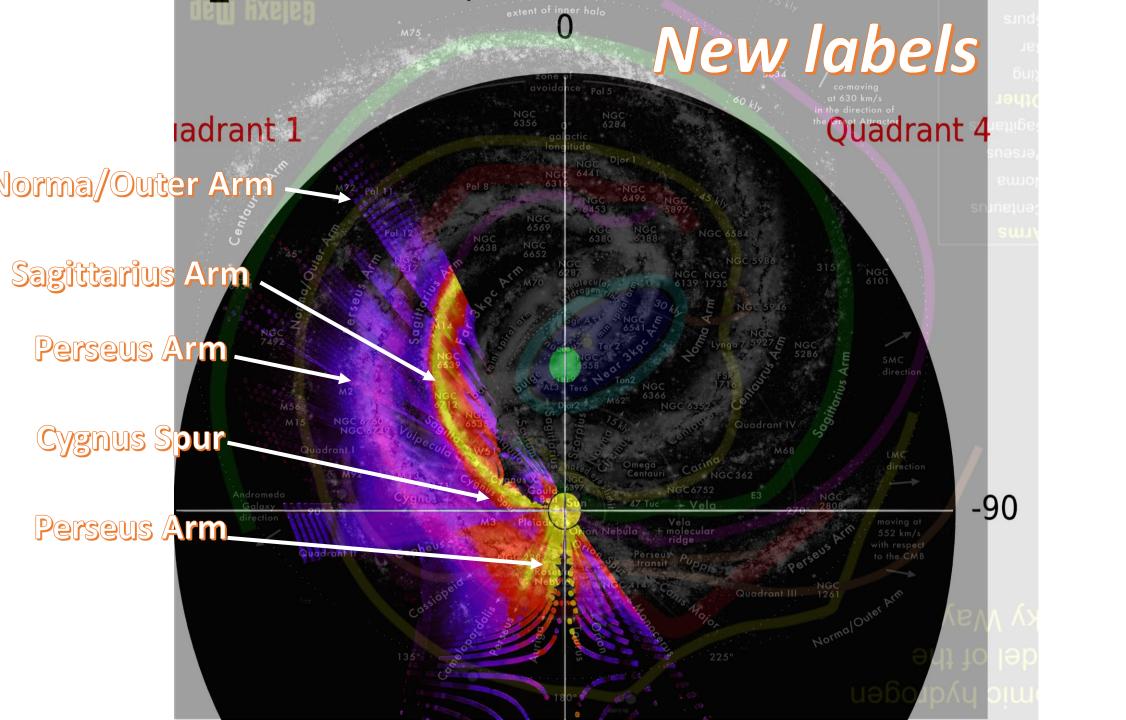




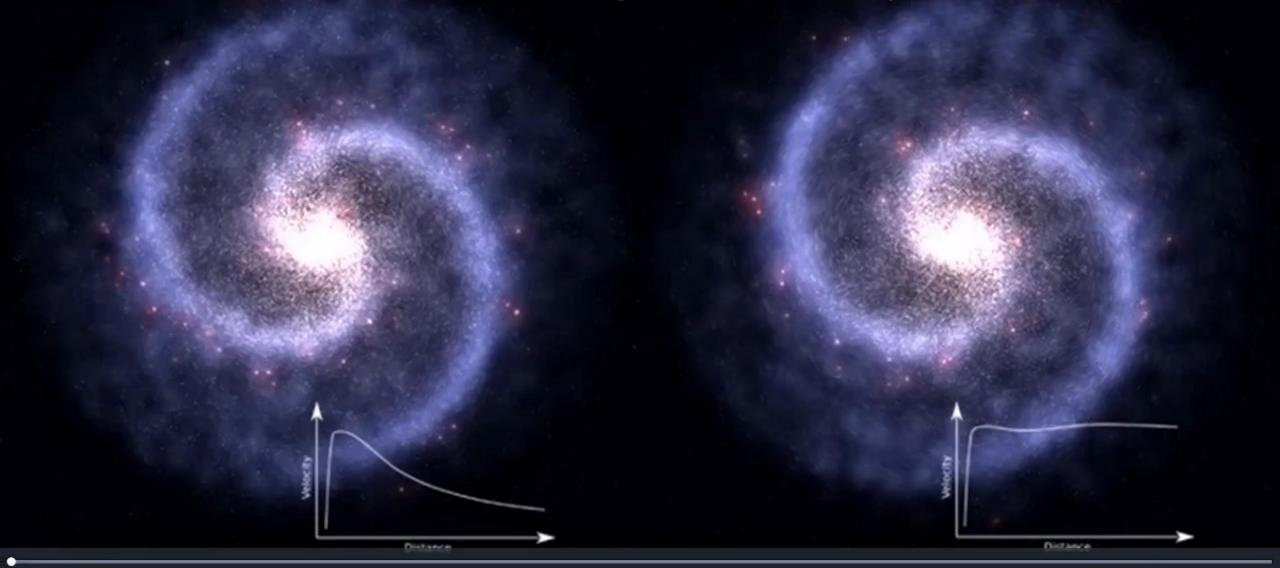
180 and -180 Galactic Longitude







### Simulation of galactic rotation WITHOUT Dark Matter (Left) vs WITH dark matter (Right); From Wikipedia



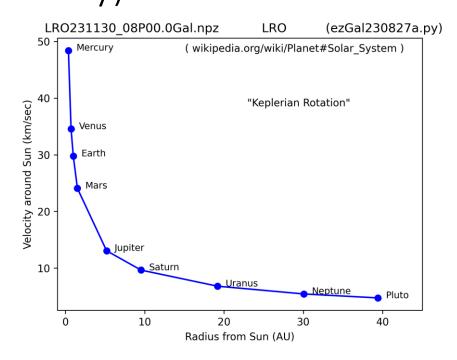


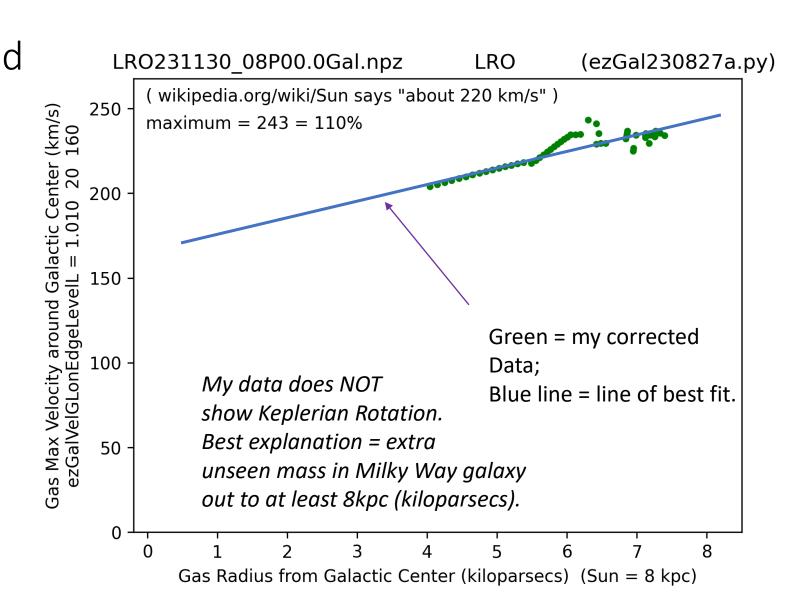




### Dark Matter - galactic rotation curve from my data

Below is typical Keplarian Rotation Curve which would be expected without dark matter, on right my data (ezRA suite/Pharmigan array)





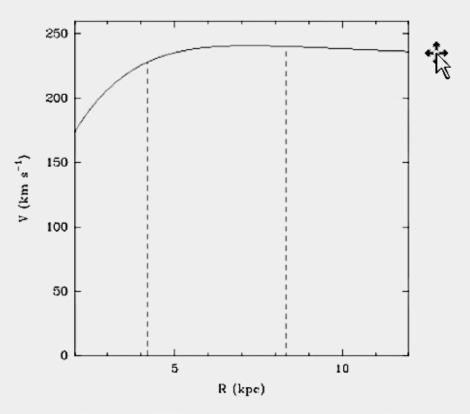
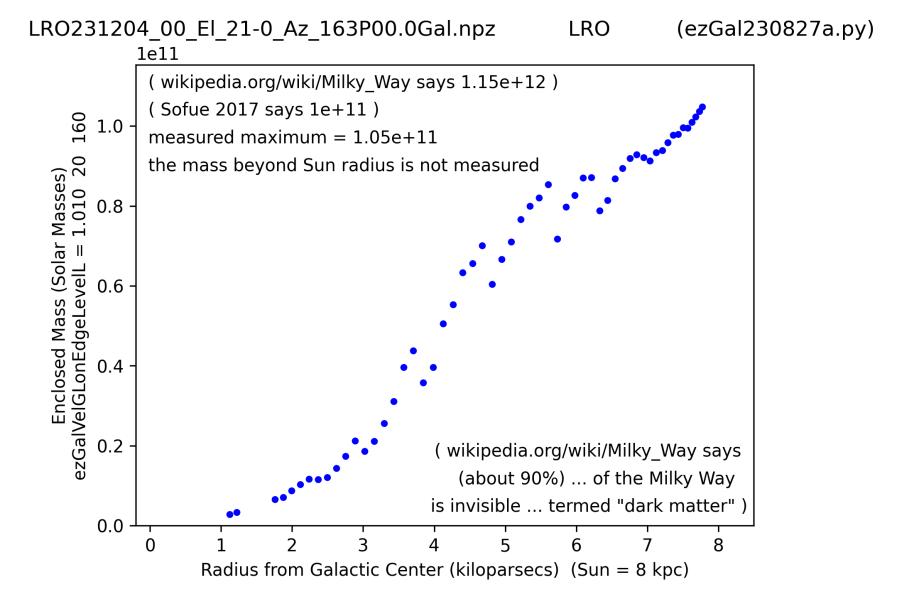


Fig. 2.— A rotation curve fitted to 80 VLBI parallaxes and proper motions of high-mass star forming regions by Reid et al. (2014), using the "Universal" rotation curve formulation of Persic, Salucci & Stell

## Compare my rotational data against this published data

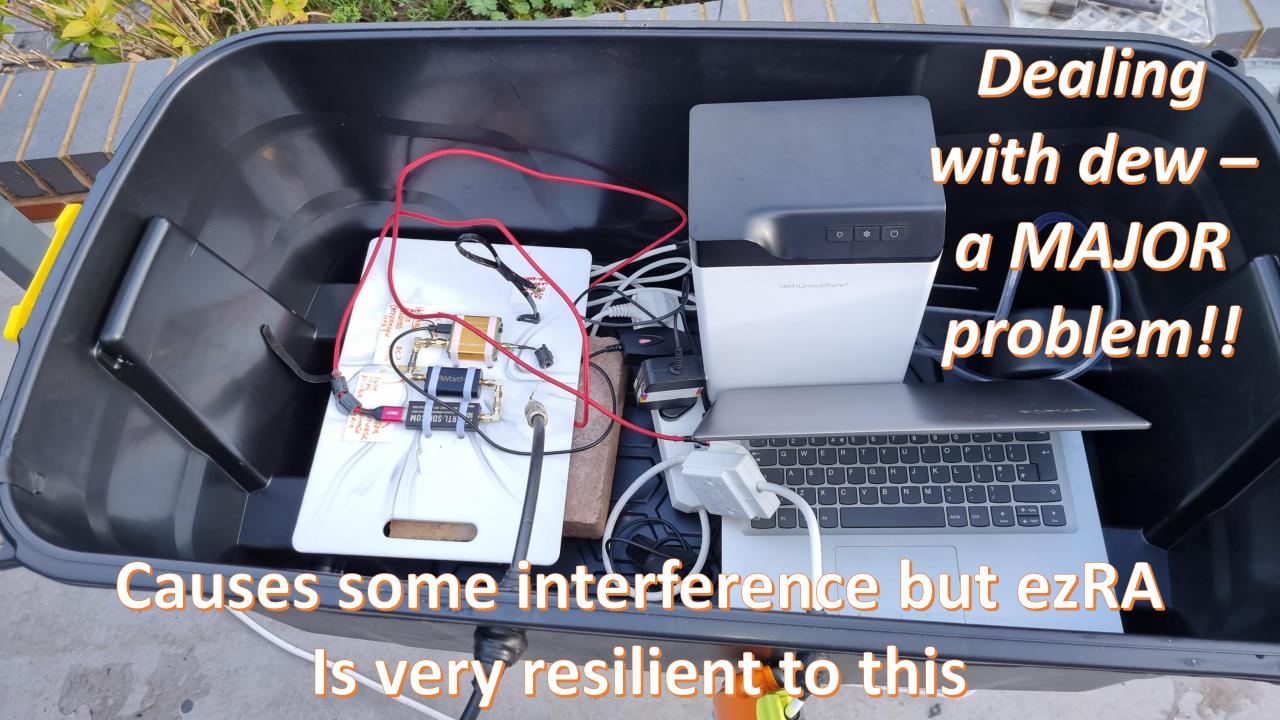
Enclosed MOSS

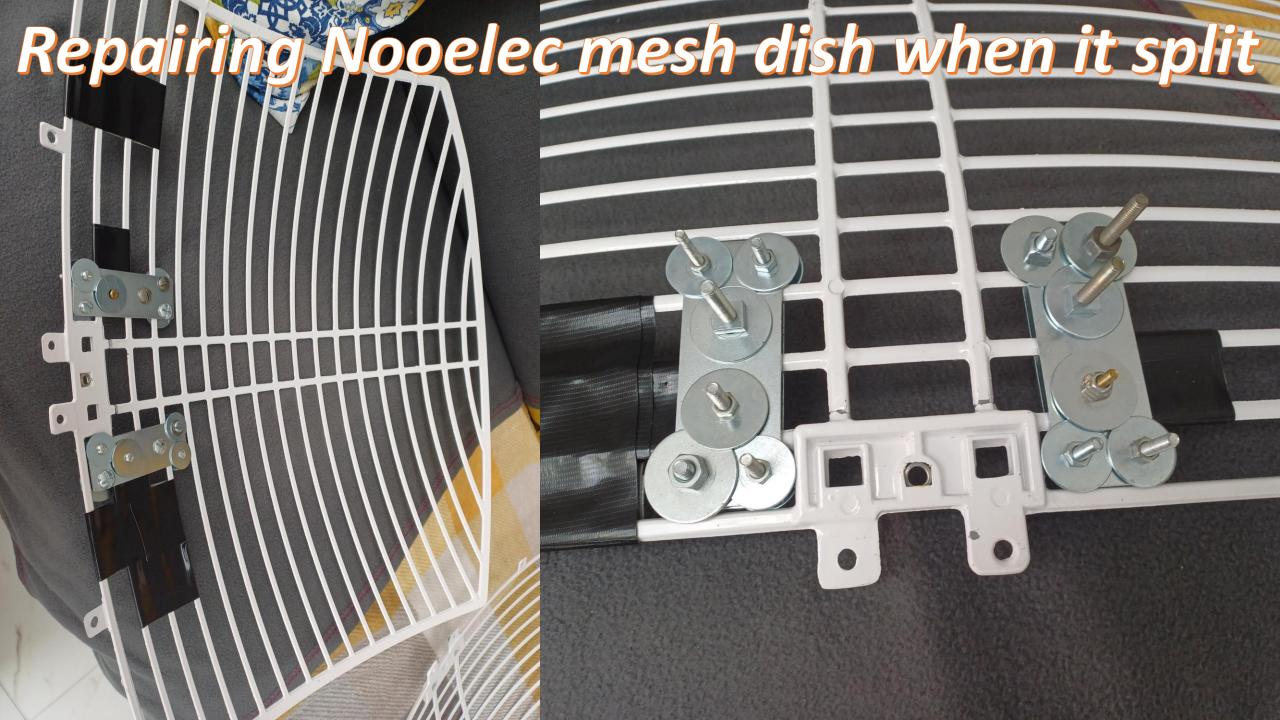


# Choosing a PC to record data– I chose Windows-based Mini-PC

- 2<sup>nd</sup> hand mini-PCs are very cheap £85 buys off ebay Levono i5, 8GB RAM, 256GB SSD. A little extra increases RAM and SSD.
- RealVNC for remote connection I opted for this rather than Remote Desktop as latter causes issues with audio inputs on remote computer – free RealVNC account gives three remote PCs per account – I now have three accounts!
- Raspberry Pis, Ardinos, Linux machines are all options too







### Contennae

My next project was
to build one of
these and try and
replicate my findings



ı	-	
ı		
ı		
ı	===:	
ı	===1	

Freq. of operation [MHz]

2450

Can Diameter [mm]

90

Cut-Off Freq. for TE11 Mode [MHz]

Cut-Off Freq. for TM01 Mode [MHz]

Waveguide Wavelength Lg =  $\lambda g$  [mm]

 $0.25 * \lambda g [mm]$ 

 $0.75 * \lambda g [mm]$ 

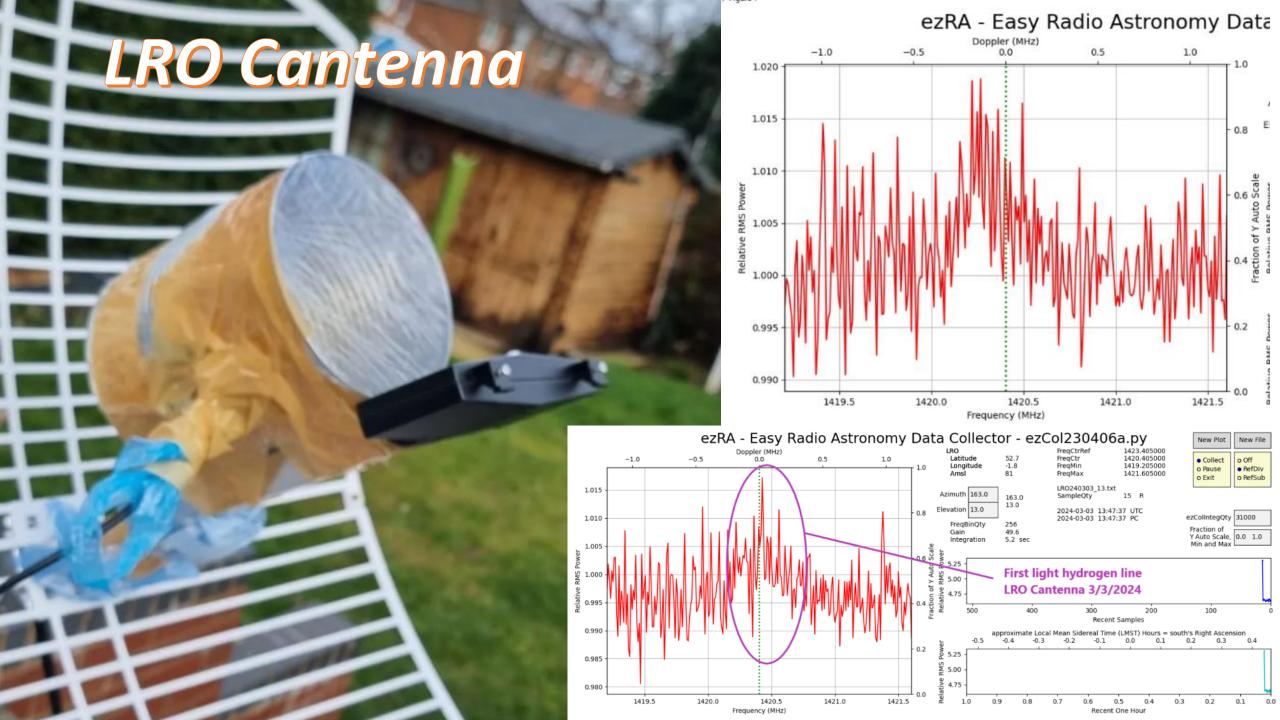
Free space Wavelength  $\lambda$  [mm]

Length of Monopole  $\lambda/4$  [mm]

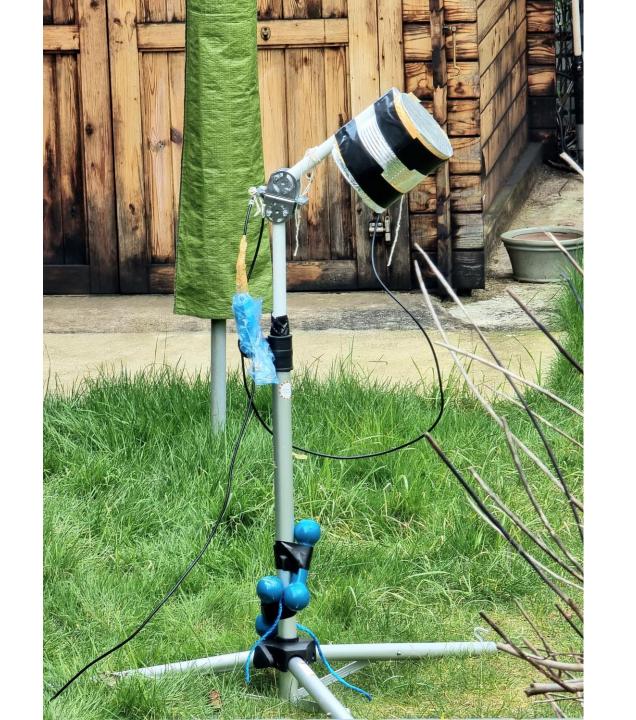
Diameter of Monopole [mm]

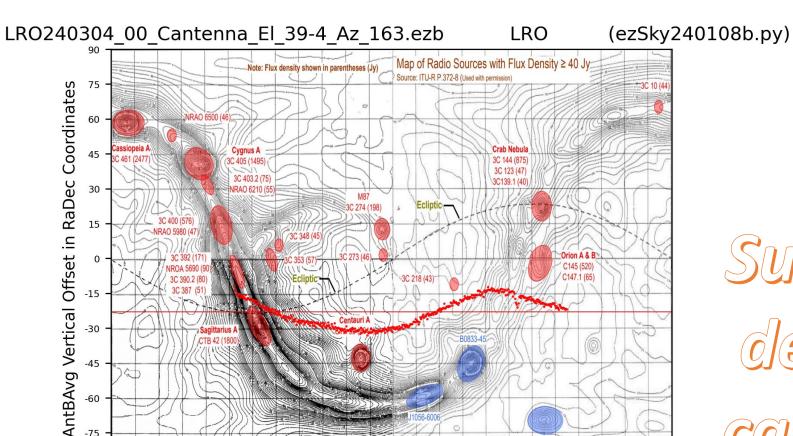
CALCULATE





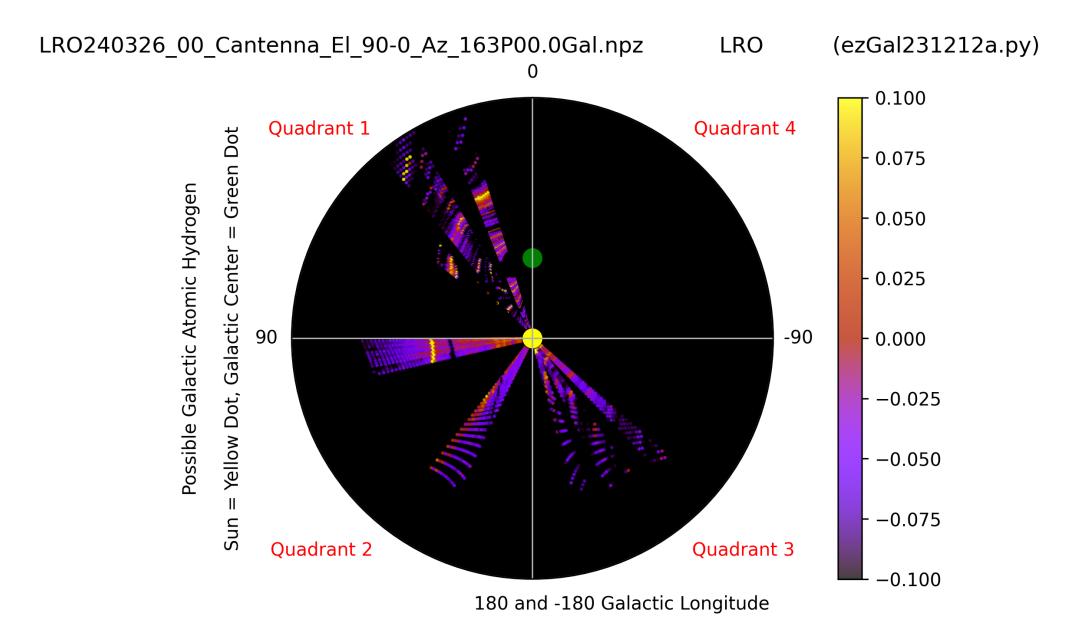






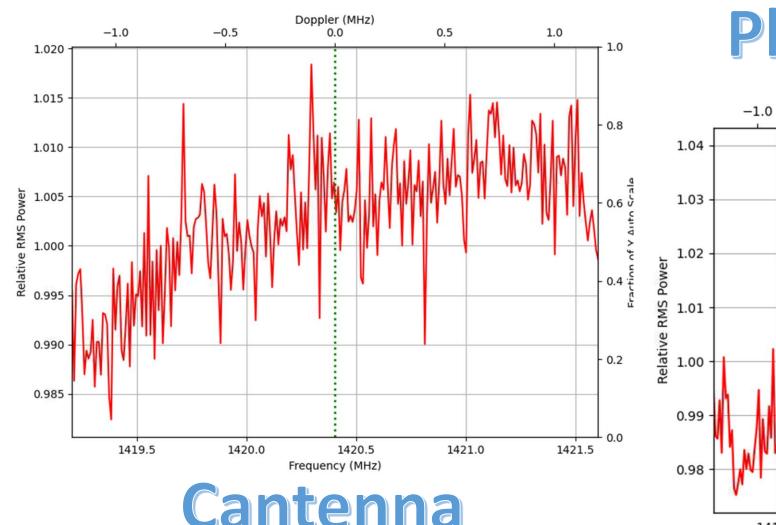
Successful h-line
detection using
contenna at LRO



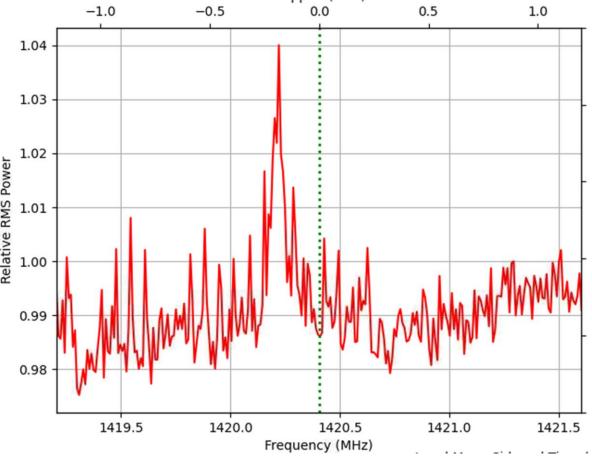


Sadly, data from LRO shows lot of interference

## Finding the problem with Cantenna Both ezCol images below 29/3/24 LRO El 56.5 Az 163



## Pharmigan Array



## IMPORTANT ADVICE FOR NEWBEES WITH LITTLE KNOWLEDGE LIKE ME:

- Get <u>support and advice</u> from SARA/BAA members.
- Realize you need to <u>dedicate lots of time</u> to get it working.
  - Making mistakes and starting again is part of the fun.
  - **Don't feel embarrassed** that you do not know enough.
- If your kit works then do not get too worried when someone else says should be done differently – E.g. your aerial has too much ground noise/your dish isn't deep enough or too deep.
  - Finally, <u>do NOT estimate</u> elevation and azimuth because your beamwidth seems large, MEASURE IT!
    - ezRA suite is quite robust to interference



### Windows Update Blocker v1.8





### Windows Updates Option

- Enable Updates
- Disable Updates
- Protect Services Settings

#### Service Status



Apply Now

Menu

**Contact Details:** Dr Andrew Thornett M6THO andrew@thornett.net www.astronomy.me.uk